From Westerbork to the Webb Telescope: 40⁺ years of Cosmic Star Formation & Supermassive Blackhole Growth

Rogier Windhorst (ASU) - JWST Interdisciplinary Scientist



Symposium honoring Prof. Harry van der Laan's 80th birthday

Sterrewacht Leiden; Saturday, October 8, 2016



- Harry van der Laan has been a truly outstanding mentor and teacher!
- He was "Chief Executive Officer" of Dutch astronomy for over a decade.
- He had the vision to involve NL in La Palma & Hawaii (see Jan's talk).
- He laid the foundation of the successes of ESO/VLT (see Tim's talk).
- Harry was always there for you when you personally needed him.



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- Harry's lectures were great, once one learned how to read them:

Syllabus from Harry's 1978 course in Extragalactic High Energy Astrophysics:

780116 Extragalactische Hogo Energie-Astropysica - NORMAN & Vanden LAAN Overzicht _ 1° holft. Datum Onderwerp Docent 6 jan. 1. Inleiding - bodens - over andt 2. Strakingsmeehanismen N = 3 jon. 1. Surveys in wrschillende spietraal geboden ; woornaam low groot heden. 6. 2. Parameter Wrivations : from absar . vables to physical parto me tons N "o jan. 122 Rodinstituts on guosons: de ampirisade stand Non rolan 6 & fot 122 Theoretical problems for R tos and 12s N 13 feb.] 122 Actin galony modei N 20 feb. Jiez incl. Sayferts 27 fab. 102 Examples: problems: rodds & ands

Extragol. High Ennyy Astrophysics Port II NURMANE Von W. LAAN Date: 15 maart Rollans 5 april Pasa 12 " 17 mei 24 mei and r werpen Clasters of golinis 3 X Dotoilal treatment of / X Do torilar treatment of on action golony 1 x Description and discussion of rodin - Xroy, WSRT- NEPO.B 1X program. Final session 1x

Harry's handwriting was not easy to read: we called it "Harry-oglyphics". Harry wrote in big "Rooster-feet" or "Hane-poten"! But you learned to Fourier transform and digest it. We learned a lot!

Westerbork traced Cosmic Star Formation and Actively Galactic Nuclei:

- Young Objects and Old objects with redshift, or
- Stellar Birth and Stellar Death over cosmic time.

Normalized differential 1.41 GHz source counts (Windhorst et al. 1985, 1993, 2003; Hopkins et al. 2000) from 100 Jy to 100 nJy. Filled circles below 10μ Jy show the 12-hr SKA simulation of Hopkins et al. (2000).

Models: giant ellipticals (dot-dash) and quasars dominate the counts to 1 mJy, starbursts (dashed) below 1 mJy. Normal spirals at cosmological distances (dot-long dash) will dominate the SKA counts below 100 nJy.

Harry whispered into my ear in June 1984: "Rogier, get involved into HST!" HST, and JWST, changed the career of this radio astronomer ...

(2a) WFC3: Hubble's new Panchromatic High-Throughput Camera

HST WFC3 and its IR channel: a critical pathfinder for JWST science.

WFC3 30 Dor: Massive stars (8-30 M_{\odot}) leave modest blackholes (3-12 M_{\odot}).

Waves that happen in Nature — Sounds Waves:

In solids: Earthquakes

In liquids: Surf!

In gasses: Sound

Electromagnetic WavesIn space-time: Gravity WavesSept. 2015: LIGO added Gravity Waves as a new way to observe Nature!

Conclusion 1: Most low-mass blackholes today are small, slow eaters:

- LIGO's 29–36 M_{\odot} blackholes: leftover from First Stars (first 500 Myr)?
- Too massive to be leftover from ordinary Supernova explosions.
- Why only seen *now* as merging by LIGO (12.5 Gyr after Big Bang)?
- They were likely not fast & efficient eaters, but slow and messy ...

JWST Hardware Status

Much progress has been made in OTE integration

Where we were at last month's call

<u>Current</u>: all 18 PMSAs installed, liquid-shim-cured, & metrologized. Alignments meet specifications, and actuator motions verified <u>Big milestone!</u>

8 February 2016 JWST Monthly Telecon 8

JWST lifetime: Requirement: 5 yrs; Goal: 10 yrs; Propellant: 14 yrs.

April 2016: NASA team-work to take JWST mirror covers off!

May 2016: JWST being tilted into the right position

May 2016: Webb mirrors finally mounted and ready!

Centaurus A NGC 5128 HST WFC3/UVIS

F225W+F336W+F438W

F502N [O III] F547M y F657N Hα+[N II] F673N [S II] F814W I

3000 light-years 1400 parsecs

56″

[TOP]: X-ray–Radio–Optical images of Cen A. [BOTTOM]: WFC3: Jet-induced SF in 2-Myr starclusters (Crockett⁺ 2009). [RIGHT]: Hydro models of bowshock-induced SF (Gardner⁺ astro-ph/1610).

-22

-24

-26

-28 -30

• Quasars: Centers of galaxies with feeding supermassive blackholes:

Hubble IR-images of the most luminous Quasars known in the universe:

- Seen at redshift $z\simeq 6$ (universe 7× smaller than today), 900 Myr old!
- Contains 10¹⁴ solar luminosities within a region as small as Pluto's orbit!
- Feeding monster blackholes (>3×10⁹ solar mass) \sim 900 Myr after BB!

Yet, the dusty(!) host galaxies are not yet visible (Mechtley $^+$ 2012, 2016).

- Who came first: Chicken or Egg?: The supermassive blackhole!
- JWST will detect 10–100× fainter dusty hosts (for $z \lesssim 20$, $\lambda \lesssim 28 \mu$ m).

Conclusion 2: Supermassive blackholes started early & were very rapid eaters:

• All massive galaxies today contain a central super-massive blackhole.

- Masses 3×10^9 solar, leftover from the First Stars (first 500 Myr)?
- Must have fed enormously rapidly in the first 1 Gyr after the Big Bang.
- Were eating *cat*-astrophically (and secretly) until they ran out of food ...
- JWST will detect 10–100× fainter dusty hosts (for z \lesssim 20, λ \lesssim 28 μ m).

Very first stars likely born in the first 500 Myr after the Big Bang.

- They were likely 80–200 solar masses, lived fast, & died young (1 Myrs!)
- They could have left 30–80 solar mass blackholes behind, as LIGO saw. JWST will observe these First Light sources after 2018:
- Expected to be weekly clustered: faint signal in JWST IR background.

The HST-unique part for JWST: Panchromatic 13 filter HUDF: UV-Blue emphasized.

592^{*h*} HUDF weighted log-log: FuvNuvUBViIzYJWH, AB \lesssim 28–31 (\gtrsim 2 nJy).

Panchromatic 13 filter HUDF.

of else-color "Balametric" or χ^2 unlige.

6

841 orbits = 592^k HUDF AB 31 mag, Objects affect ~45% of pixelsU

HST Frontier Field A2744: JWST needs lensing to see First Light at $z\gtrsim10-15$.

Conclusions: JWST First Light surveys must consider three aspects: (1) The very rapid drop in space density (LF) for $z\gtrsim 8$.

(2) Cannot-see-the-forest-for-the-trees effect ["Natural Confusion" limit]: Background objects blend into foreground because of their own diameter.

(3) House-of-mirrors effect ["Gravitational Confusion"]:

• JWST needs to find most First Light objects at $z\gtrsim 10-15$ through the best cosmic lenses (making the images even more crowded):

• Lensing is needed to see what Einstein thought was impossible to observe!

Thank you, Harry, for everything you did for us:

• For your excellent mentorship!

• For your career-changing advise!

• For your friendship and understanding!

• And for constantly supporting us!

Integrated UV–IR galaxy counts converge to yield accurate Extragalactic Background Light estimates (Driver⁺ ApJ, 827, 108; astro-ph/1605.01523). • Integrated starlight and dust ($z_{med} \simeq 1.5$) each contribute ~50%. Direct γ -ray Blazar measurements may suggest dim excess at $\lambda \sim 1-5\mu$ m: • JWST will constrain *diffuse* Pop III star component at $z\simeq 10-20$.

Spitzer 3–5 μ m power-spectrum with galaxies removed (Kashlinsky⁺ 2012).

- JWST's superior spatial resolution will substantially improve discrete galaxy light subtraction:
- JWST can detect any diffuse Pop III star excess at $\lambda \simeq 1-5 \mu$ m.
- JWST will constrain direct-collapse or primordial blackhole models (Kashlinsky 2016).

Will this ever happen to our own Galaxy?

YES! Hubble showed no lateral motion: Approaches at -110 km/s. Hence, Andromeda will merge with Milky Way!

The two blackholes (10^6-10^7 suns) will also merge!

Not to worry: only 4–5 Gyr from today!

Illustration Sequence of the Milky Way and Andromeda Galaxy Colliding

NASA, ESA, Z. Levay and R. van der Marel (STScI), T. Hallas, and A. Mellinger - STScI-PRC12-20b

SuperMassive BlackHole mass vs. Galaxy Bulge Mass

(For elliptical galaxies only)

0.5% of total galaxy mass makes it into SMBH!

SMBH=cosmic garbage disposal: Messy leftover of galaxy formation!

(Kormendy & Ho, 2013 An Rev A&Ap 51, 511)