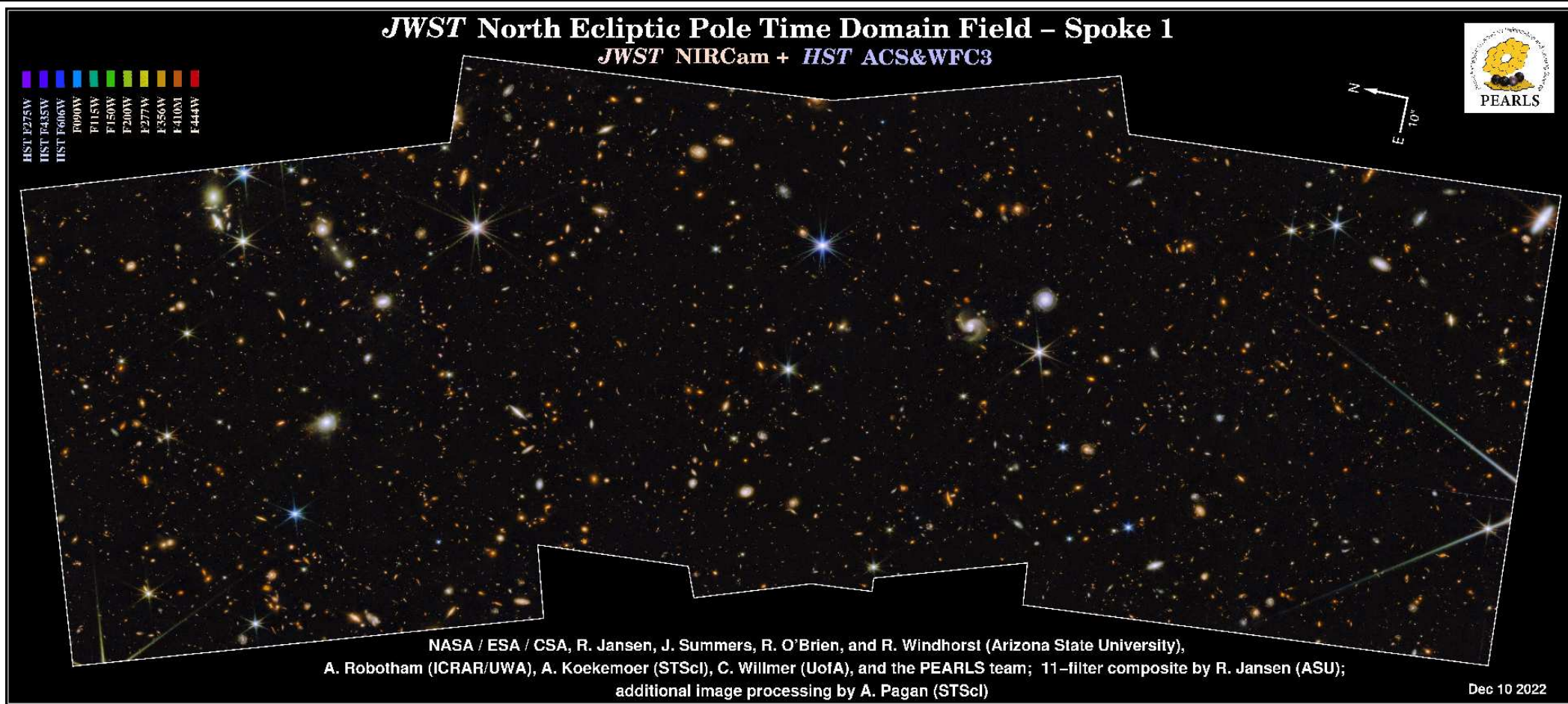


# The Crown Jewels of the JWST PEARLS Project

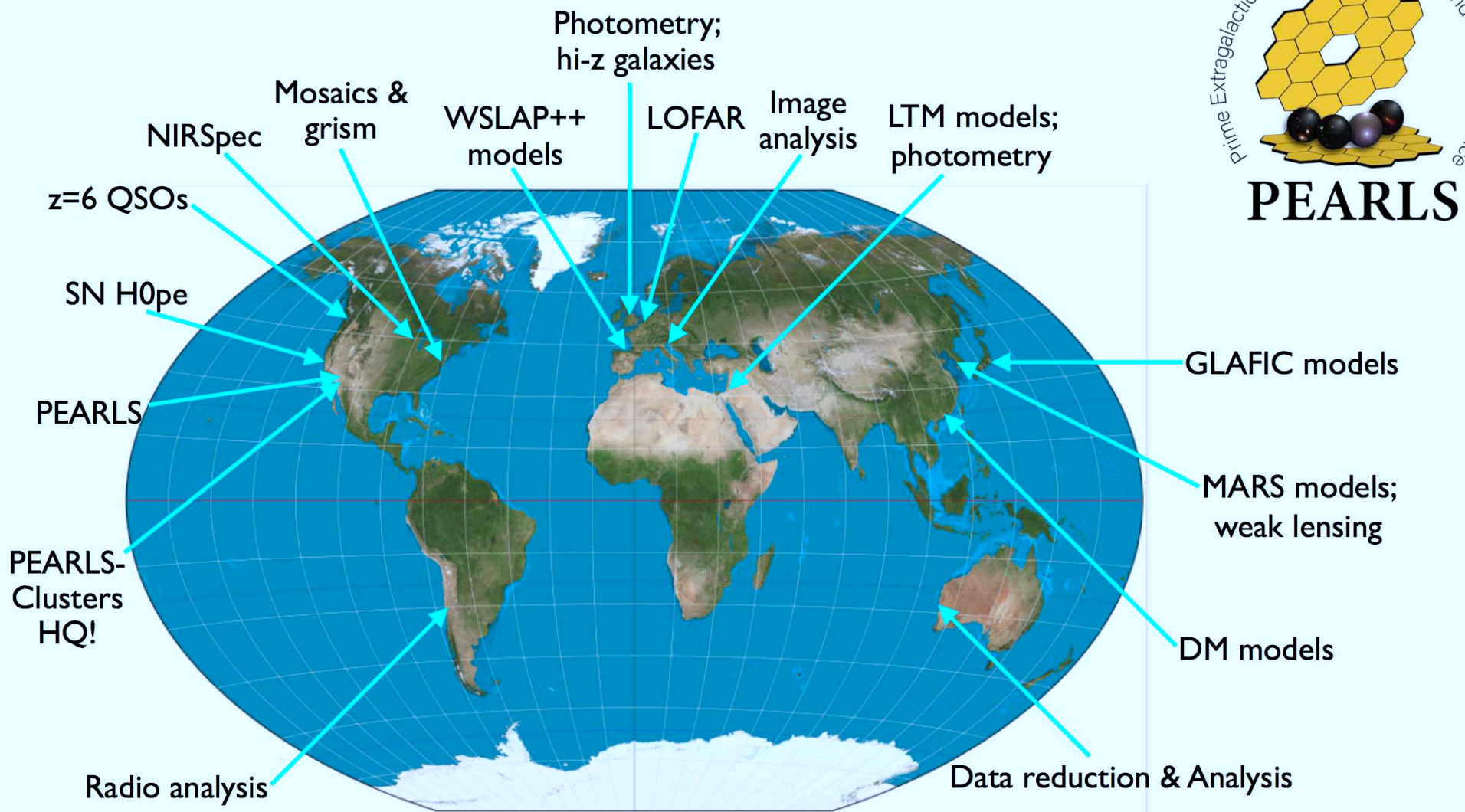
Rogier Windhorst (ASU) — Regents' Professor & JWST Interdisciplinary Scientist

+JWST PEARLS team: T. Carleton, S. Cohen, R. Jansen, P. Kamienieski, T. Acharya, J. Berkheimer, N. Foo, R. Honor, D. Kramer, T. McCabe, I. McIntyre, R. O'Brien, R. Ortiz, J. Summers, C. Conselice, J. Diego, S. Driver, J. D'Silva, B. Frye, A. Koekemoer, M. Marshall, M. Nonino, A. Robotham, S. Tompkins, C. Willmer, H. Yan, N. Adams, D. Austin, D. Carter, D. Coe, K. Duncan, H. Hammel, N. Grogin, W. Keel, N. Pirzkal, M. Polletta, R. Ryan Jr., I. Smail, S. Willner, R. Arendt, J. Beacom, R. Bhatawdekar, L. Bradley, T. Broadhurst, C. Cheng, F. Civano, L. Dai, H. Dole, G. Fazio, G. Ferrami, L. Ferreira, S. Finkelstein, L. Furtak, H. Gim, A. Griffiths, K. Harrington, N. Hathi, B. Holwerda, J. Huang, M. Hyun, M. Im, B. Joshi, I. Juodzbališ, P. Kelly, R. Larson, J. Li, J. Lim, Z. Ma, P. Maksym, G. Manzoni, A. Meena, S. Milam, M. Pascale, A. Petric, A. Pozo Laroche, P. Porto, C. Redshaw, H. Rottgering, M. Rutkowski, S. Scheller, B. Smith, A. Straughn, L. Strolger, A. Swirbul, J. Trussler, L. Wang, B. Welch, S. Wilkins, S. Wyithe, M. Yun, E. Zackrisson, J. Zhang & X. Zhao et al. (120 scientists over 18 time-zones)



The First Year of JWST Science Conference; Monday Sept. 11, 2023 (STScI, Baltimore, MD)

# PEARLS Program



PEARLS = Prime Extragalactic Areas for Reionization and Lensing Science (Windhorst<sup>+</sup> 2023, AJ, 165, 13):

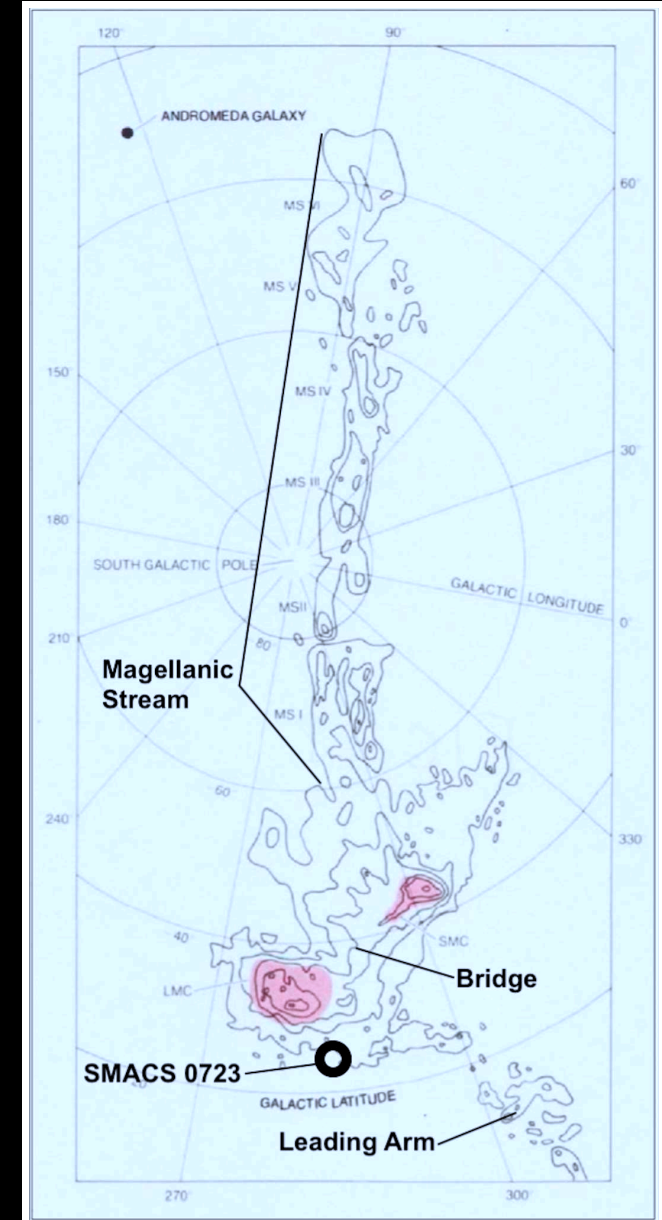
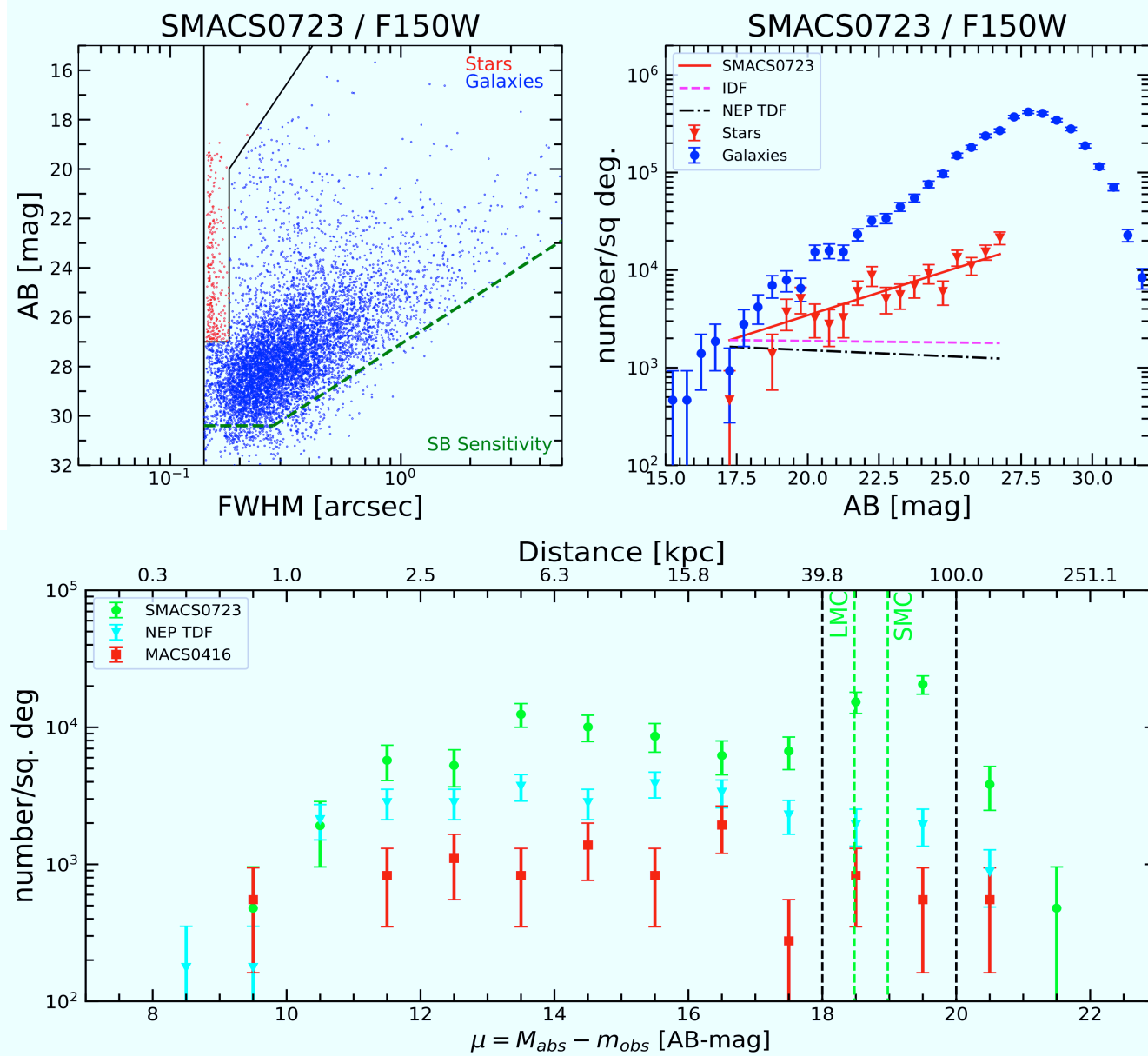
- A mix of medium-deep NIRCcam fields (GTO-2738; PIs Windhorst & Hammel), best lensing clusters (GTO-1176 Windhorst & DD-4446 PI Frye), and high-zs QSOs (GTO-1176 & GO-1813 PI Marshall).
- PEARLS crown jewels today: Extremes in Cosmic SF (low  $\rightarrow$  high  $\sim 10^7 \times$ )!



North Ecliptic Pole (NEP) Time Domain Field (TDF) from PEARLS project

— some remarkable results in PEARLS and other JWST projects:

- (Old star) tidal tails everywhere (J. Summers<sup>+</sup> astro-ph/2306.13037);
- $\lesssim 1\%$  of objects variable: AGN & SNe (R. O'Brien, R. Jansen<sup>+</sup> 2023);
- Gravitational (galaxy-galaxy) lensing common (Keel<sup>+</sup> 23, AJ, 165, 166).

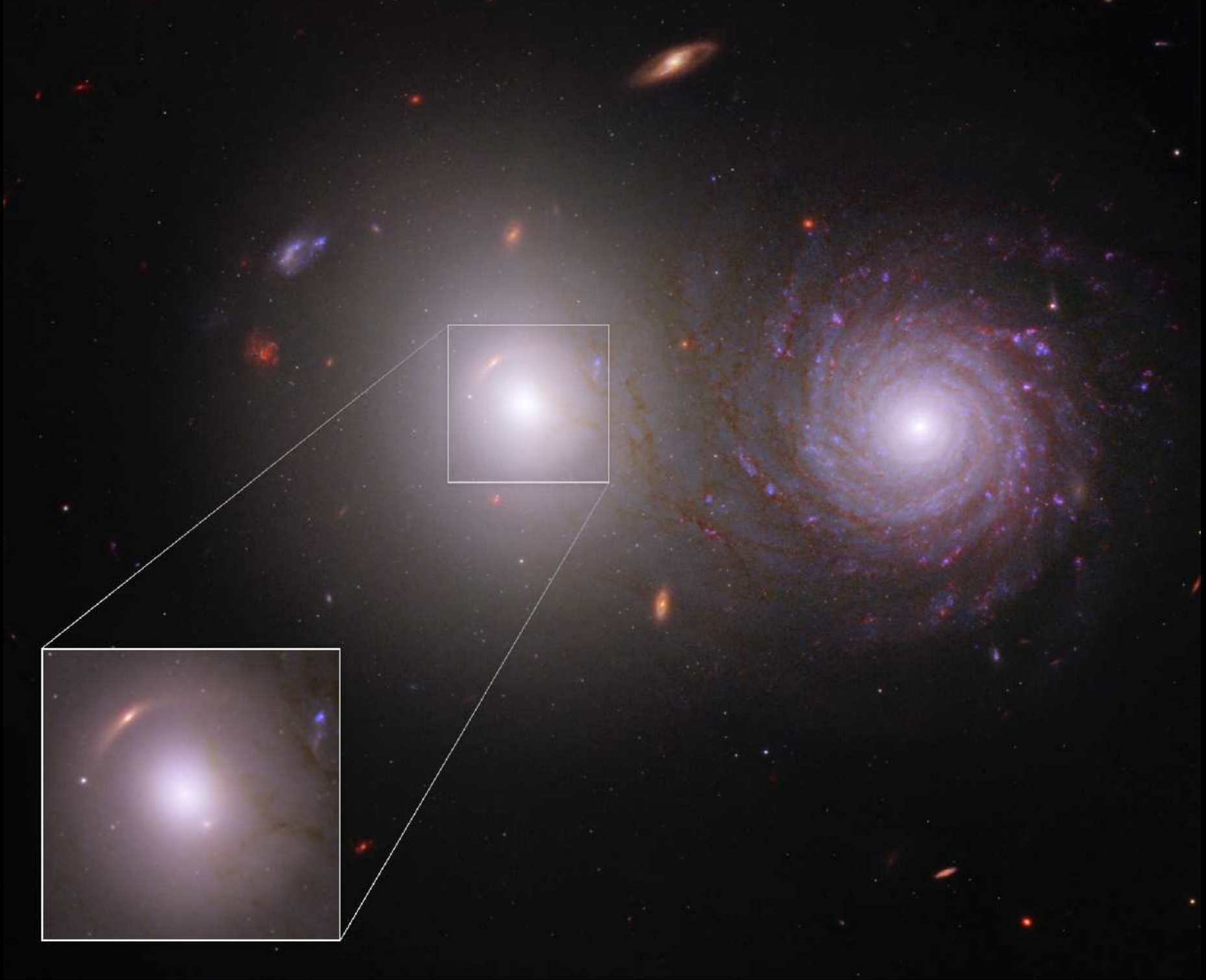


Summers, J., et al. (astro-ph/2306.13037):

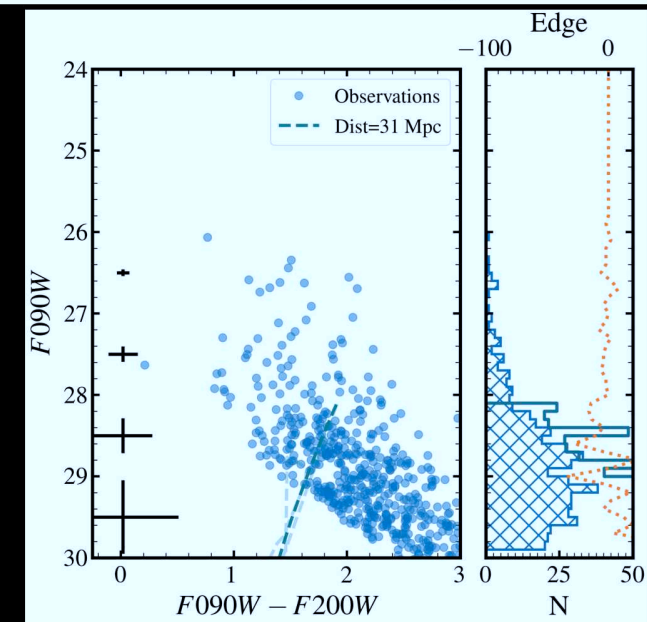
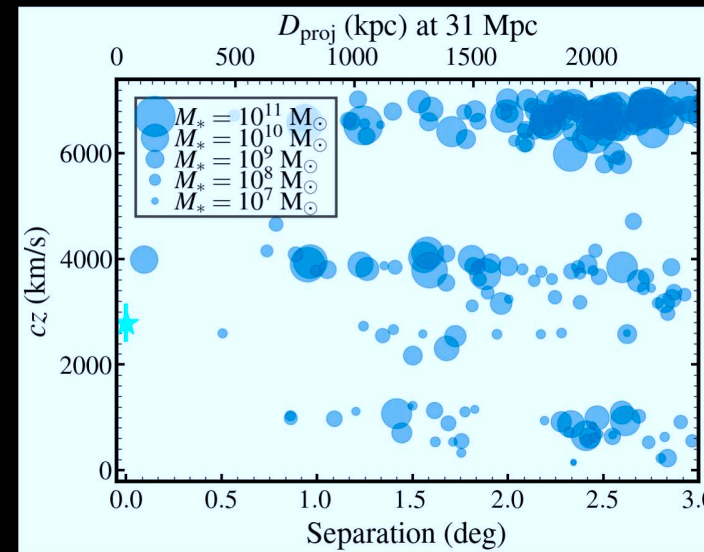
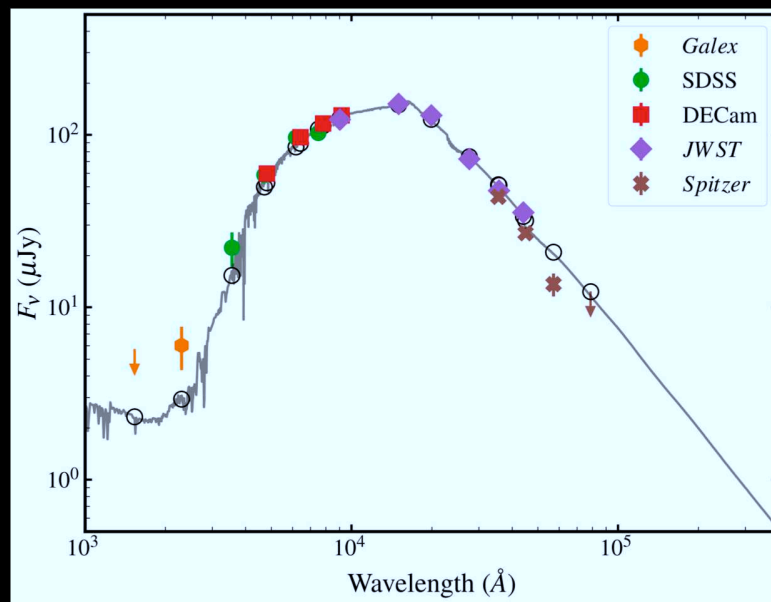
- SMACS0723 star counts show excess near LMC/SMC compared to NEP;
- 71 stars ( $AB \lesssim 27$ ) with  $D_{SpecType} \simeq 40-100$  kpc about  $10^\circ$  from LMC;
- Part of Leading Arm between LMC and MW: 10 mag fainter than Gaia!



- Spiral overlapping Elliptical: Trace cosmic dust: small grains! (Keel<sup>+</sup> 2023).
- 100's of Globular Clusters in Elliptical at  $z=0.0513$  (J. Berkheimer<sup>+</sup> 2023).



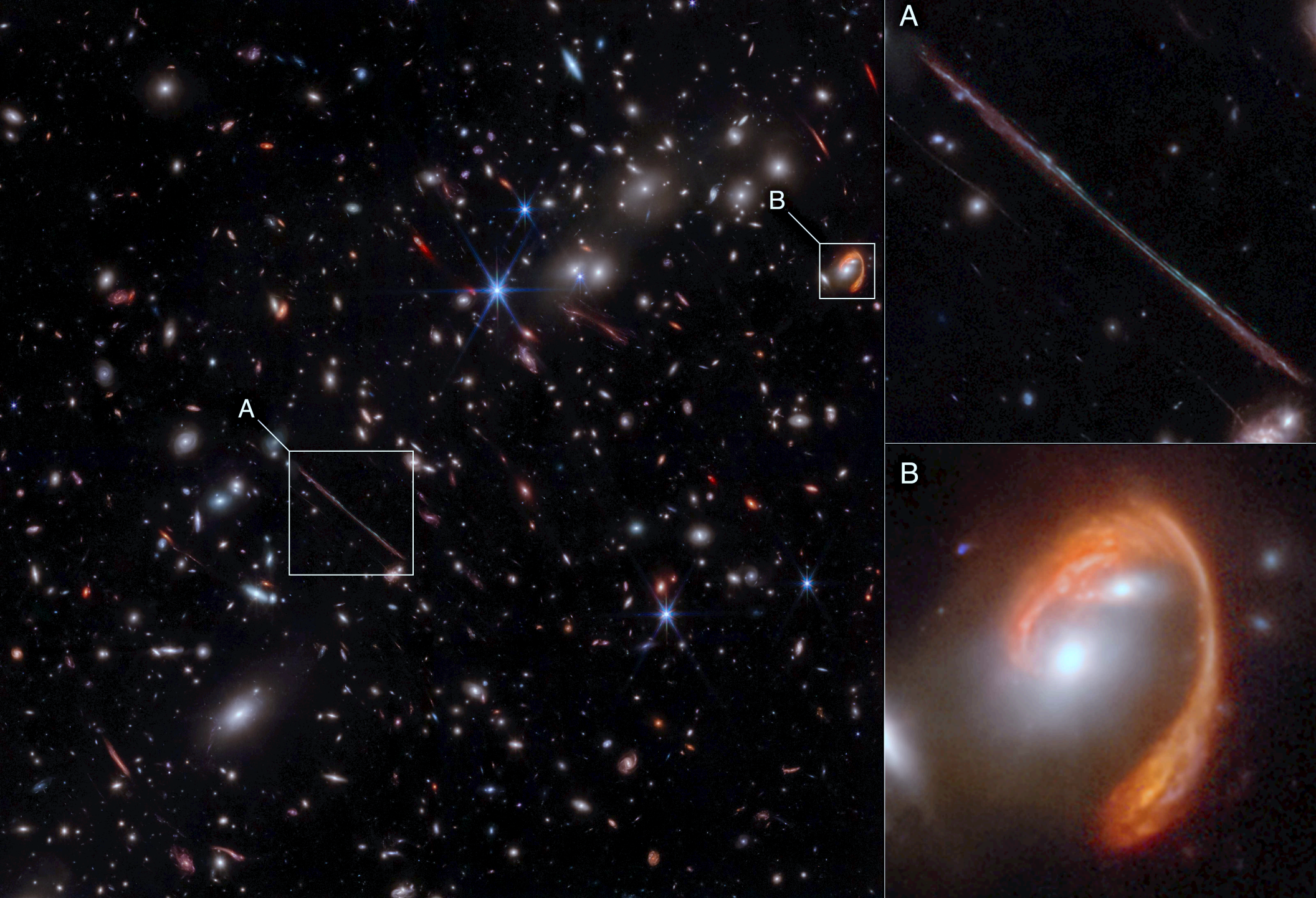
... and the  $z=0.0513$  Elliptical also lenses a background galaxy at  $z\sim 1$  (Keel, et al. 2023, AJ, 165, 16)!



PEARLSDG - an isolated quiescent dwarf galaxy  
with a TRGB distance of 31 Mpc

The Carleton<sup>+</sup> 2023 NIRC*am* image of  
A1489 (LTM cluster CLG1212) revealed:

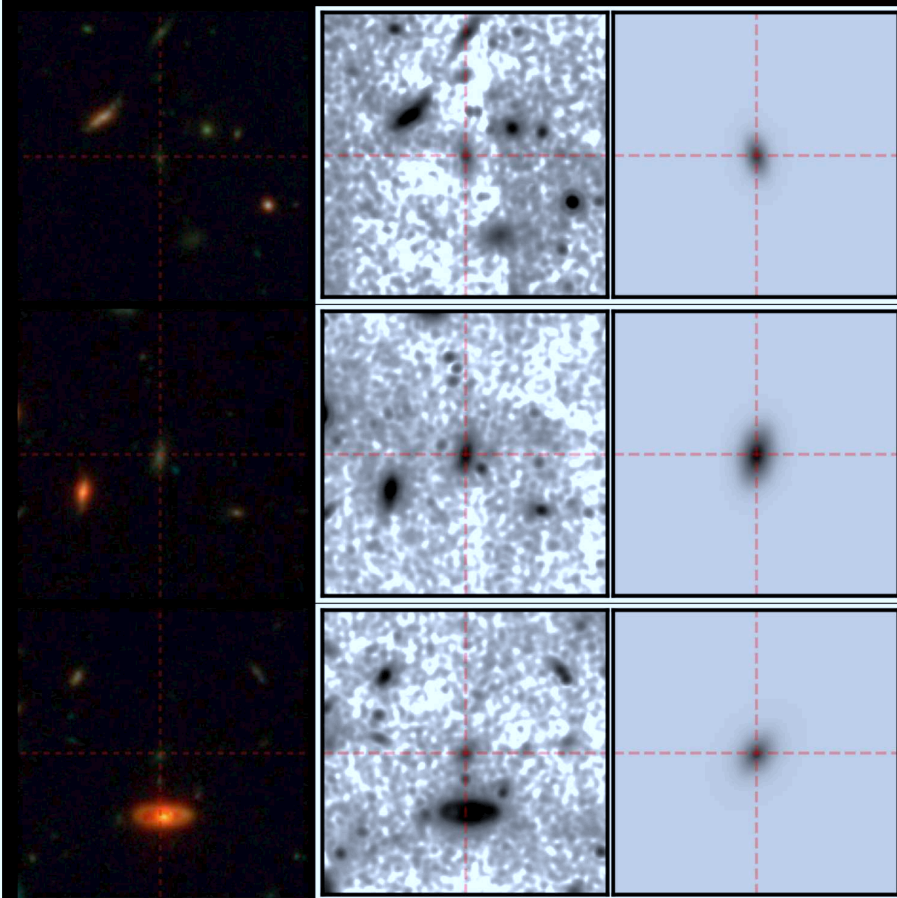
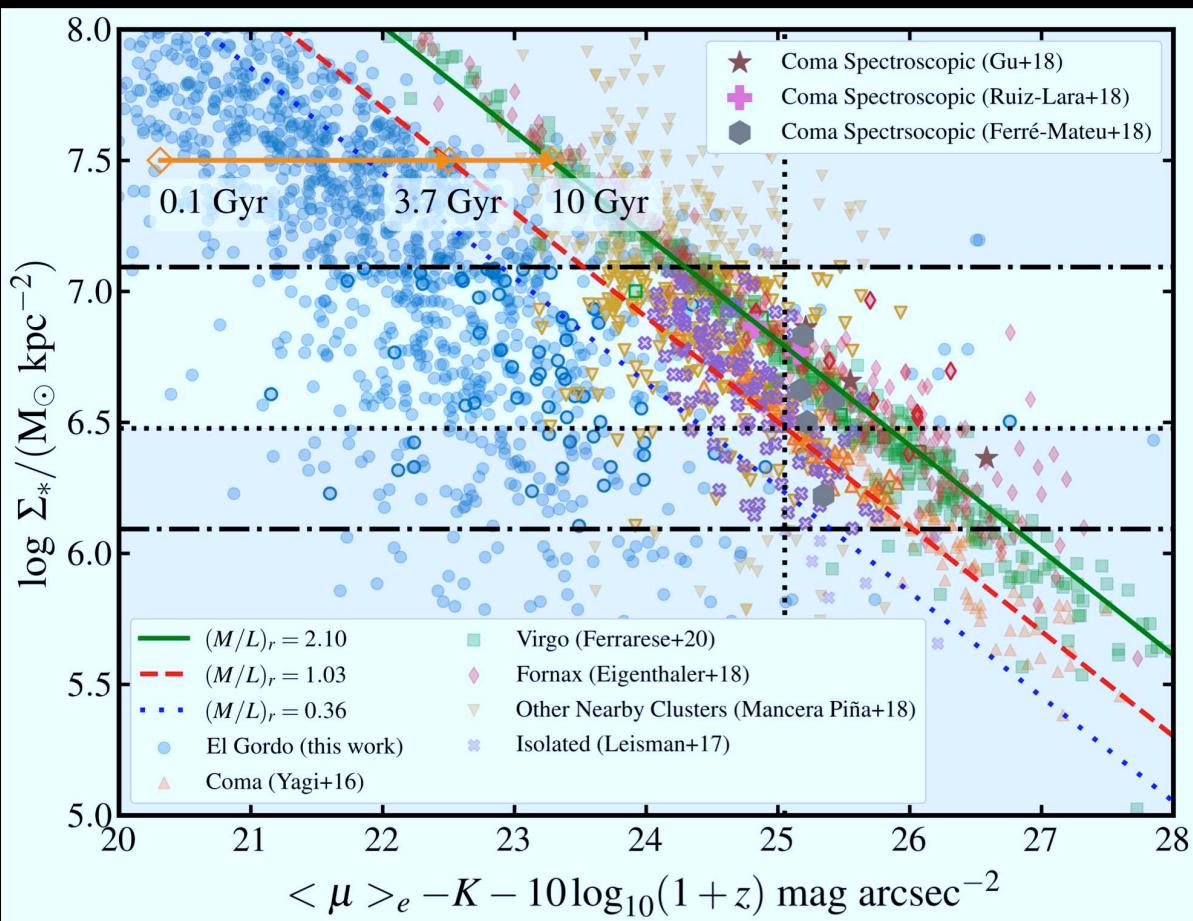
- Isolated Low-SB dwarf at  $z_{spec}=0.010$  between Virgo and Coma LSS;
- TRGB distance of  $31.0 \pm 1.6$  Mpc,  $\gtrsim 800$  km/s from nearest neighbor;
- Rather red & old, with extremely low SFR  $\simeq 10^{-4} M_{\odot}/\text{yr}$ !



## 8-filter JWST/NIRCam of massive El Gordo cluster at redshift $z \approx 0.87$

T. Carleton<sup>+</sup> (2023, ApJ, 953, 83); P. Kamieneski<sup>+</sup> 23 (astro-ph/2303.05054); J. Diego<sup>+</sup> (2023; A&A, 672, A3); B. Frye, N. Foo<sup>+</sup> (2023, ApJ, 952, 81).





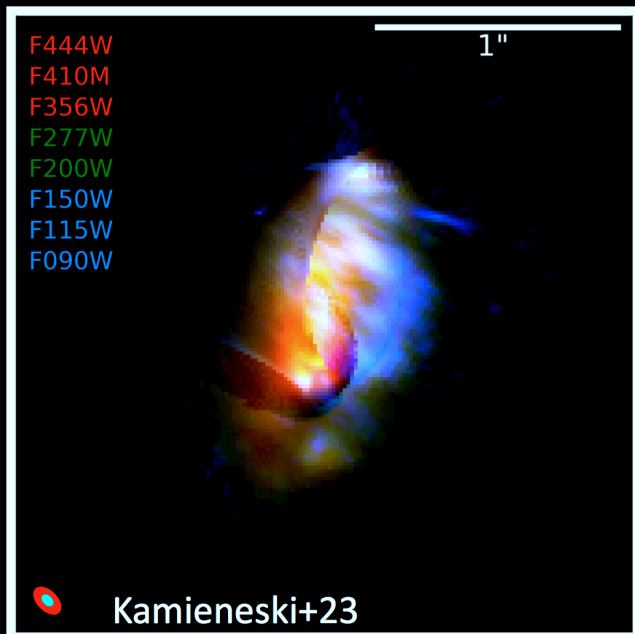
We find low stellar density galaxies in imaging of the El-Gordo cluster.  
 These objects are younger than local UDGs, suggesting that they were more recently accreted onto the cluster.

Carleton<sup>+</sup> (2023, ApJ, 953, 83; astro-ph/2205.06347) NIRCam:

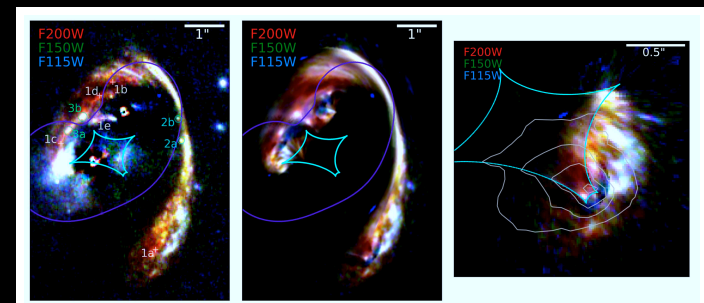
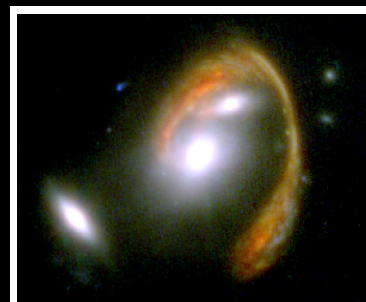
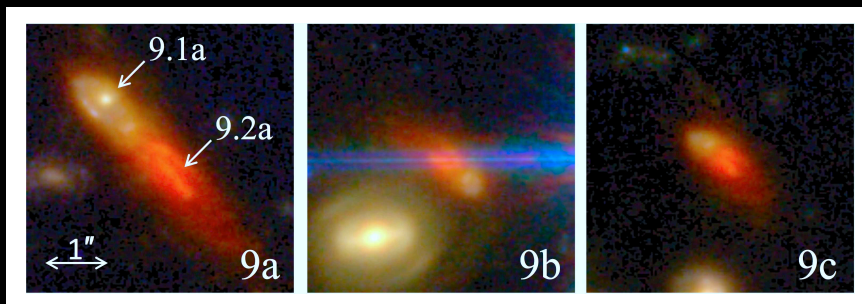
- El Gordo's Low Density Galaxies at  $z=0.87$  have  $\text{SFR} \simeq 0.1 M_{\odot}/\text{yr}$ .
- *i.e.*, SFR higher than Local Dwarfs, but of course much lower than the upcoming higher- $z$  extremes ...

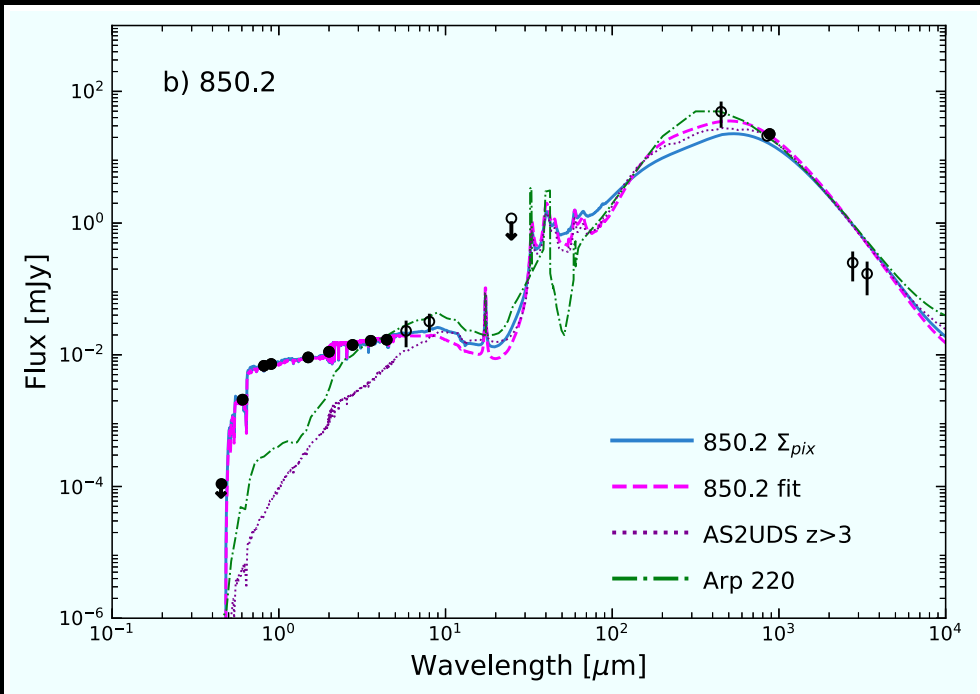
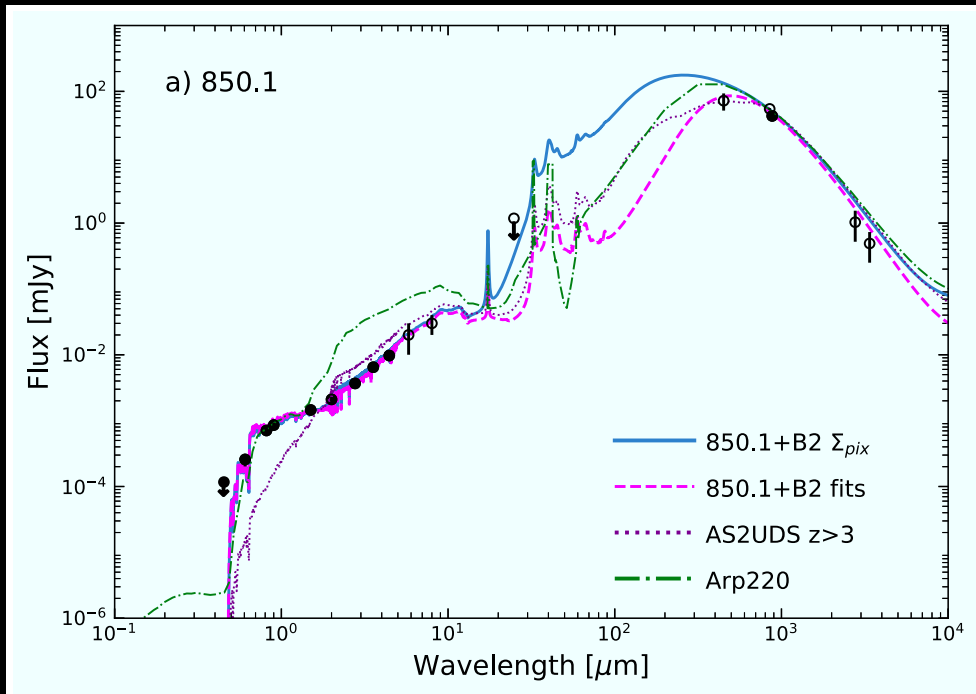
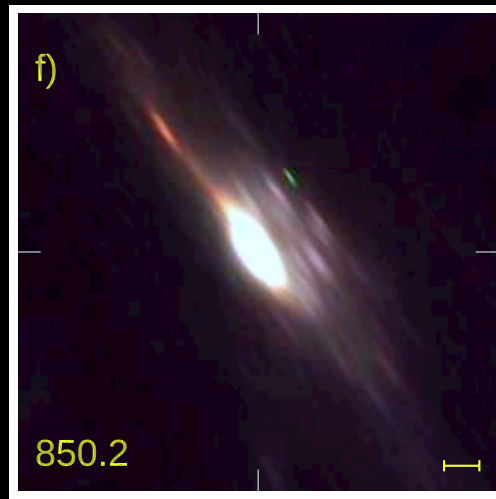
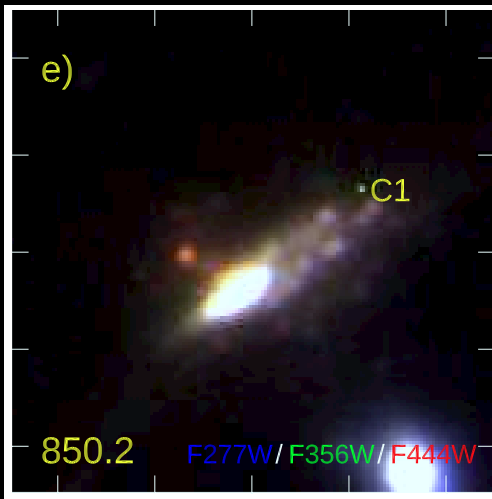
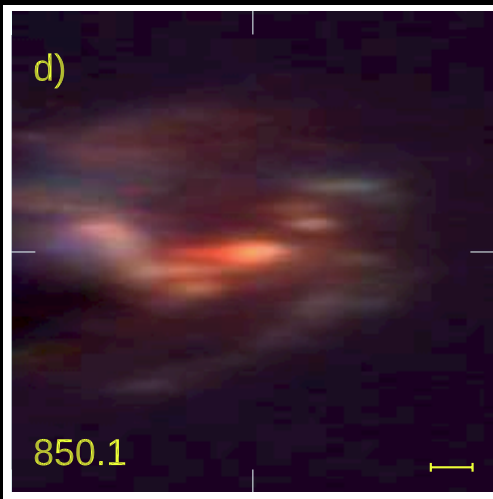
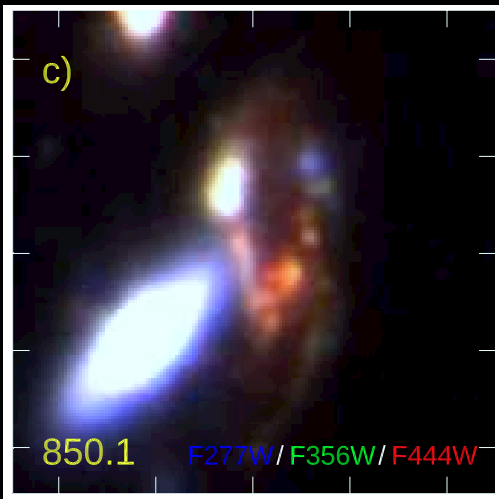
# El Anzuelo “The Fish Hook”

- SFR  $\sim 80 M_{\odot}/\text{yr}$
- $A_V \sim 2$
- Source plane:



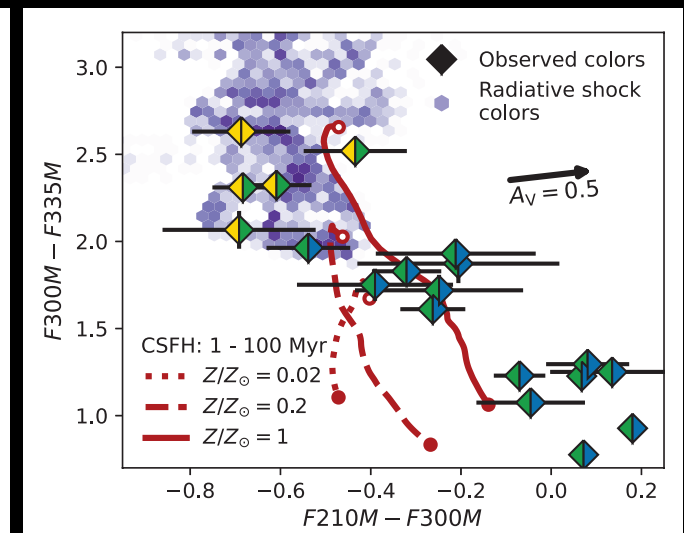
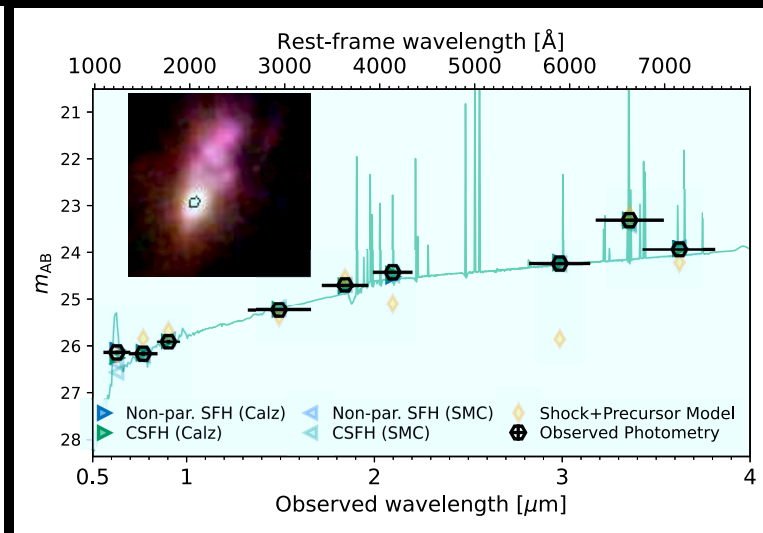
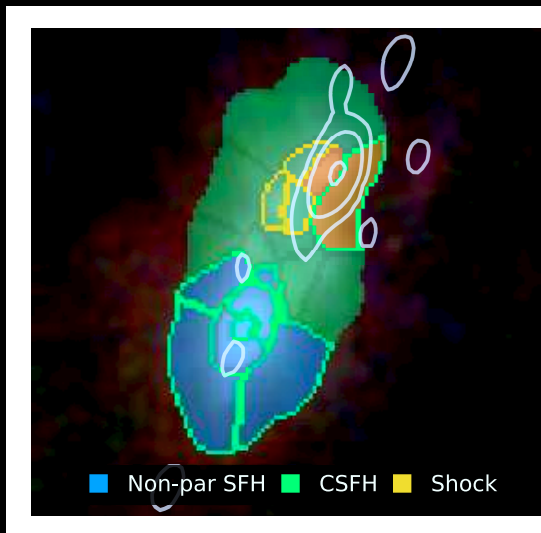
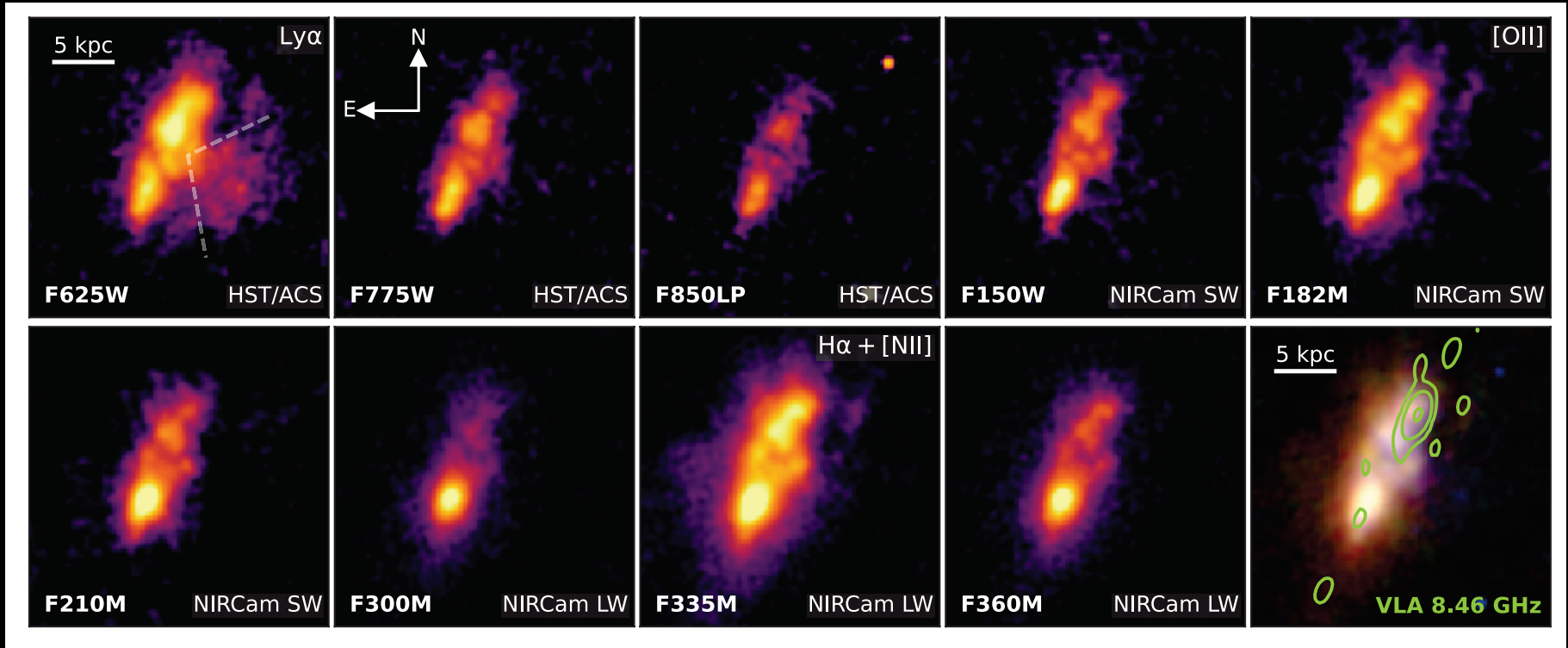
Dusty “El Anzuelo” has a high de-magnified total SFR  $\simeq 80 M_{\odot}/\text{yr}$  (P. Kamieneski<sup>+</sup> 23; astro-ph/2303.05054):





Smal<sup>+</sup> (astro-ph/2306.16039): Two bright lensed sub-mm galaxies at  $z \simeq 4.26$  behind A1489 (LTM cluster CLG1212) could not be more different:

- 850.1:  $\sim 10^{11.8} M_{\odot}$ ,  $\tau \sim 450$  Myr,  $1400 M_{\odot}/\text{yr}$ ,  $A_V \sim 5$  mag!
- 850.2:  $\sim 10^{10.3} M_{\odot}$ ,  $\tau \sim 50$  Myr,  $400 M_{\odot}/\text{yr}$ ,  $A_V \sim 1.2$  mag!

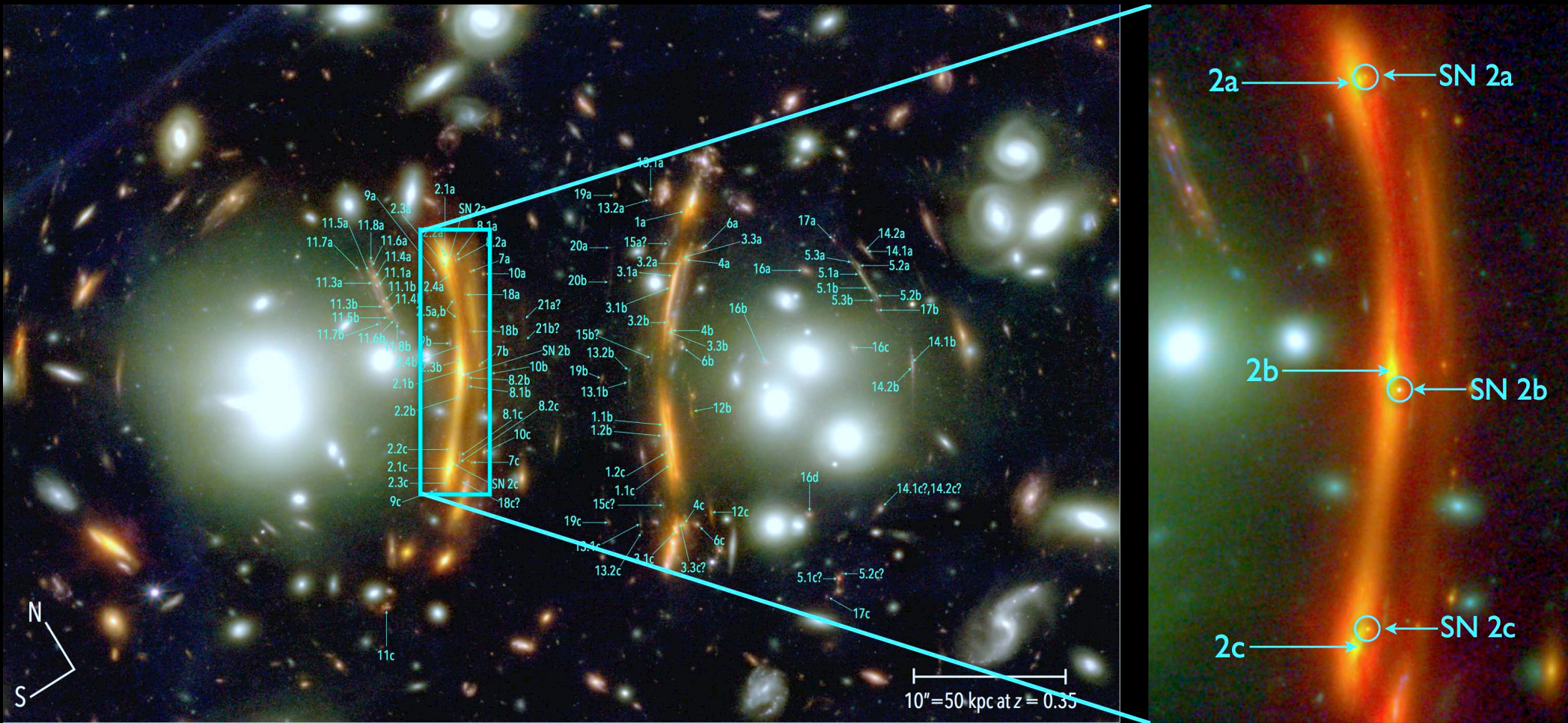


The most massive ( $10^{10.9} M_{\odot}$ ) high- $z$  radio galaxy TNJ1338 at  $z=4.11$ :  
 Total medium-band  $SFR \simeq 1600 M_{\odot}/yr$  (Duncan<sup>+</sup>23, MNRAS, 522, 4548)  
 • Extreme jet-induced  $SFR \gtrsim 500 M_{\odot}/yr$  and  $t_{SFR} \simeq 4$  Myr.



NIRCam image of most luminous far-IR Planck cluster G165 at  $z=0.35$ :

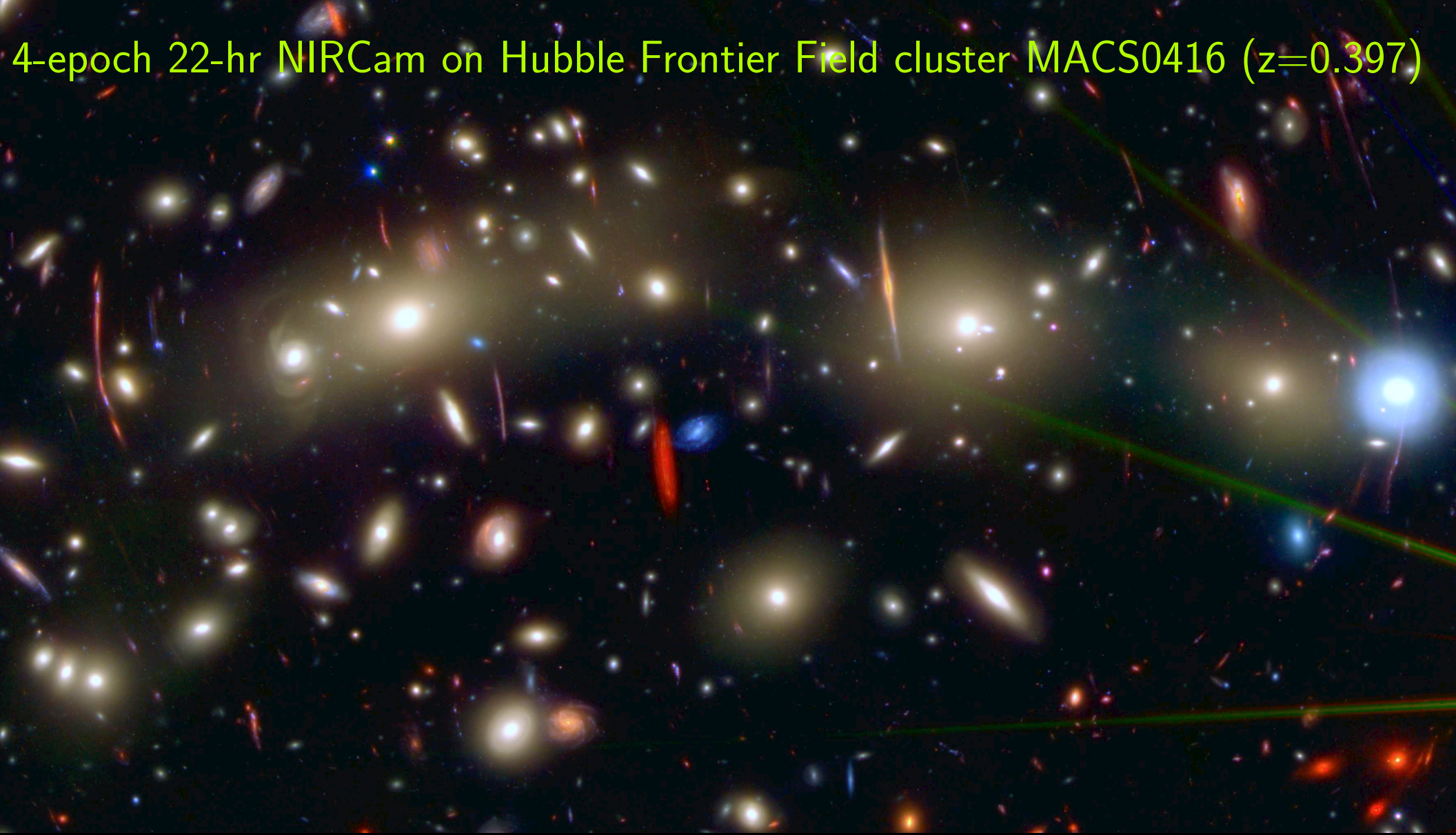
- Frye<sup>+</sup> 23: Very high *de-magnified total*  $SFR \simeq 200\text{--}350 M_{\odot}/\text{yr}$ .



NIRCam in G165 shows: 3 bright point sources parity-flipped w.r.t. Arc-2:

- Clear SN-Ia at  $z=1.783$ ! (LBT, NIRSspec; Polletta<sup>+</sup>, 2023, A&AL, 675, L4; astro-ph/2306.12385);
  - 3-epoch G165: 9 data-point light-curve!  $\longrightarrow$  suitable to measure  $H_0$  at  $z=1.783$  (Frye<sup>+</sup>23; Chen<sup>+</sup>23; Foo<sup>+</sup>23; Pascale<sup>+</sup>23; Pierel<sup>+</sup>23);
- $\longrightarrow$  Regular monitoring of clusters with extreme SF can yield more lensed SNe!
- Total SFR  $\simeq 200\text{--}350 M_{\odot}/\text{yr}$  should give a lensed SN every few years.

# 4-epoch 22-hr NIRCam on Hubble Frontier Field cluster MACS0416 ( $z=0.397$ )



- Yan, H.<sup>†</sup> (2307.07579): 12 new caustic transits at  $z \simeq 1-2$  from 4 epochs!
  - Diego, J.<sup>†</sup> (2307.10363): extremely magnified  $z=2.091$  binary star!
- ⇒ Regular monitoring of several clusters can yield IMF's at  $z \gtrsim 1$  directly!

# References and other sources of material

Talk: [http://www.asu.edu/clas/hst/www/jwst/stsci23\\_JWST\\_1stYr\\_Windhorst.pdf](http://www.asu.edu/clas/hst/www/jwst/stsci23_JWST_1stYr_Windhorst.pdf) Data: <https://sites.google.com/view/jwstpearls>

- Carleton, T., Windhorst, R. A., O'Brien, R., et al. 2022, *AJ*, 164, 170 (astro-ph/2205.06347)
- Carleton, T., Cohen, S. H., Frye, B., et al. 2023, *ApJ*, 953, 83 (astro-ph/2303.04726)
- Diego, J. M., Meena, A. K., Adams, N. J., et al. 2023, *A&A*, 672, A3 (astro-ph/2210.06514)
- Diego, J. M., Sun, B., Yan, H., et al. 2023, *A&A*, in press (astro-ph/2307.10363)
- Duncan, K. J., Windhorst, R. A., et al. 2023, *MNRAS*, 522, 4548–4564 (astro-ph/2212.09769)
- Frye, B. L., Pascale, M., Foo, N., et al. 2023, *ApJ*, 952, 81 (astro-ph/2303.03556)
- Juodzbališ, I., Conselice, C. J., Singh, M., et al. 2023, *MNRAS*, 525, 1353 (astro-ph/2307.07535)
- Kamieneski, P. S., Frye, B. L., Pascale, M., et al. 2023, *ApJ*, in press (astro-ph/2303.05054)
- Keel, W. C., Windhorst, R. A., Jansen, R. A., et al. 2023, *AJ*, 165, 166 (astro-ph/2208.14475)
- Polletta, M. del Carmen, Nonino, M., Frye, B., et al. 2023, *A&AL*, 675, L4 (astro-ph/2306.12385)
- Robotham, A. S. G., D'Silva, J. C. J., Windhorst, R. A., et al. 2023, *PASP*, 135, 085003 (astro-ph/2305.01175)
- Summers, J., Windhorst, R. A., Cohen, S. H., et al. 2023, *ApJ*, resubmitted (astro-ph/2306.13037)
- Windhorst, R., Timmes, F. X., Wyithe, J. S. B., et al. 2018, *ApJS*, 234, 41 (astro-ph/1801.03584)
- Windhorst, R. A., Cohen, S. H., Jansen, R. A., et al. 2023, *AJ*, 165, 13 (astro-ph/2209.04119)
- Yan, H., Cohen, S. H., Windhorst, R. A., et al. 2023, *ApJL*, 942, L8 (astro-ph/2209.04092)
- Yan, H., Ma, Z., Sun, B., et al. 2023, *ApJ*, resubmitted (astro-ph/2307.07579)

<https://hubblesite.org/contents/news-releases/2022/news-2022-050>

<https://blogs.nasa.gov/webb/2022/10/05/webb-hubble-team-up-to-trace-interstellar-dust-within-a-galactic-pair/>

<https://blogs.nasa.gov/webb/2022/12/14/webb-glimpses-field-of-extragalactic-pearls-studded-with-galactic-diamonds/>

<https://esawebb.org/images/pearls1/zoomable/>

<https://webbtelescope.org/contents/news-releases/2023/news-2023-119>

<https://news.asu.edu/20230801-jwsts-gravitational-lens-reveals-distant-objects-behind-el-gordo-galaxy-cluster>

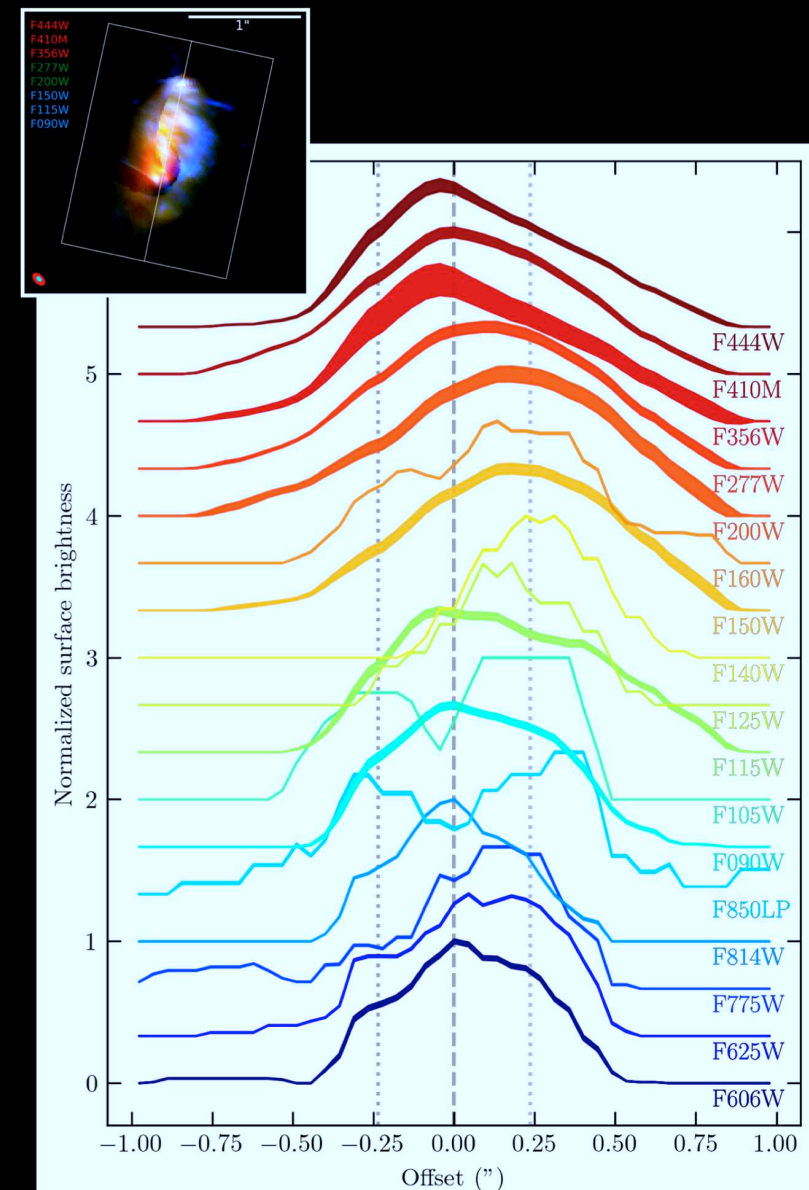
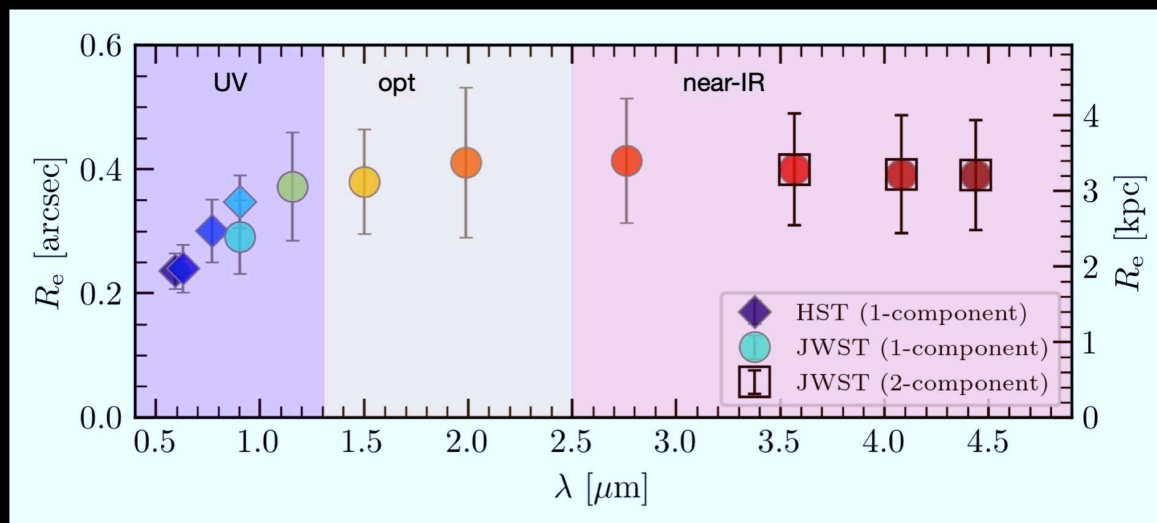


# SPARE CHARTS

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## How does its size vary from UV to IR?

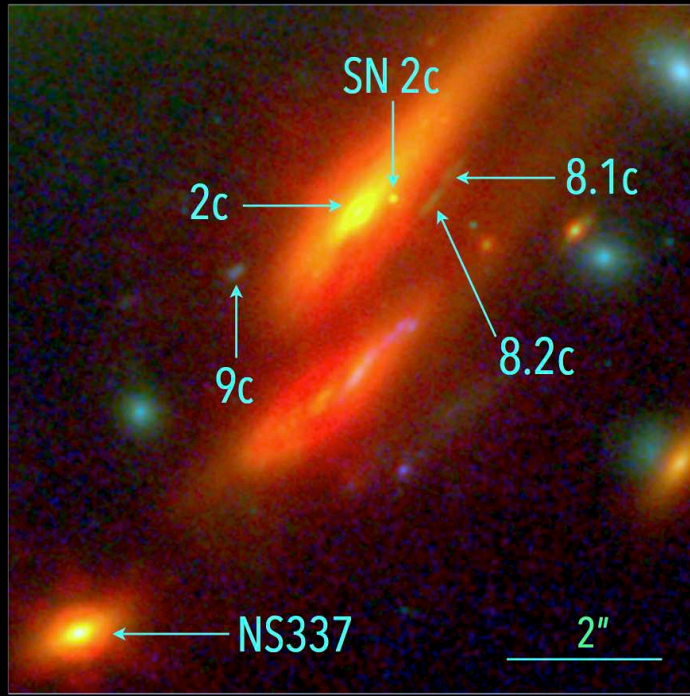
- Effective radius is nearly *constant* from rest-frame blue optical through near-IR, despite a redder center
- Larger sSFR in galaxy outskirts --> Inside-out quenching?
- Both variations in sSFR and in dust attenuation are likely responsible for the complex color gradients



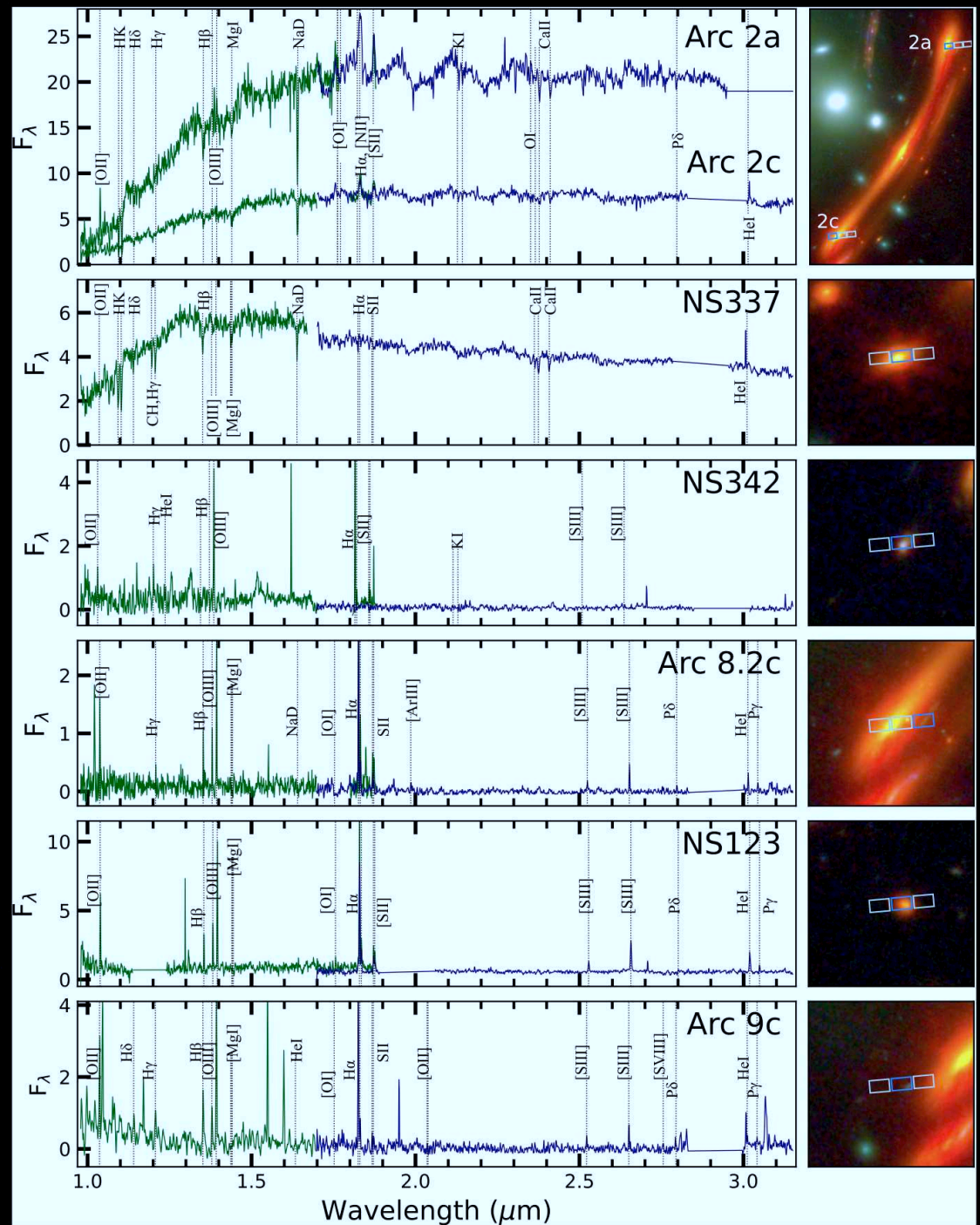
Kamieneski et al. (2023, ApJ, in press; astro-ph/2303.05054):

- Dusty “El Anzuelo” has a high de-magnified total  $\text{SFR} \simeq 80 M_{\odot}/\text{yr}$ .
- Larger sSFR in delensed outskirts  $\implies$  inside-out quenching?

# JWST/NIRSpec



- SN 2c and its host galaxy (2c) are depicted with 4 *close* galaxy neighbors ( $\Delta v < 900$  km/s, source plane separation  $< 34$  kpc)
- Spectroscopy/SED fits find Arc 2 to be quiescent & massive & its friends to be SFGs/SBGs
- This picture is potentially consistent with galaxy downsizing
- Look for upcoming papers (Frye+23b; Pierel+23; Chen+23; Pascale+23b)



● Frye<sup>+</sup> 23: Very high *de-magnified* total SFR  $\simeq 200\text{--}350 M_\odot/\text{yr}$ .